

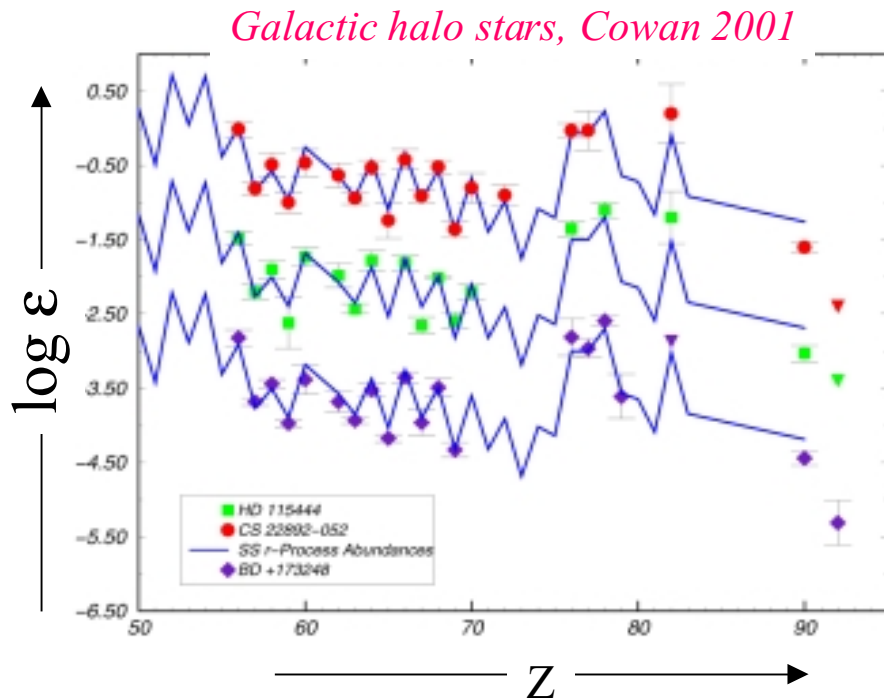
Neutron capture cross sections around neutron-rich magic nuclei for astrophysical purposes

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- Abundances observations
 - Astrophysical process
 - Experimental β -decay studies
- Neutron captures mechanisms
- Neutron captures on $^{46}\text{Ar}_{28}$

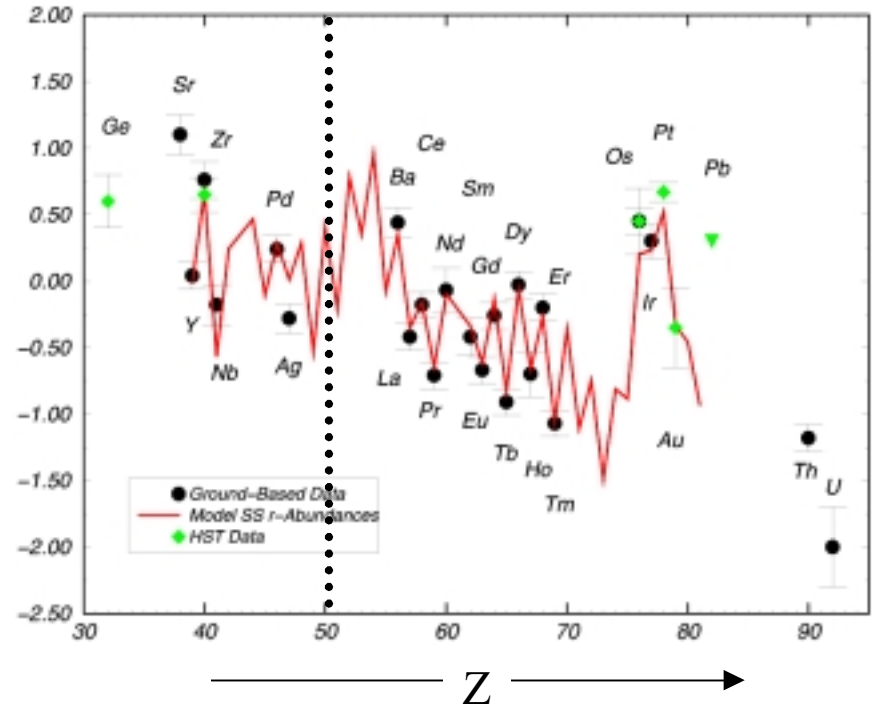
Observations in UMP Stars



r abundances **similar** to solar system
for $Z > 54$, $A > 130$

Same abundance **pattern** in **3** different
stars from the galactic halo

Very robust r -process for $A > 130$



For $Z < 54$: non solar

Weak r -process (weak neutron
fluxes) responsible for
production of $A < 130$ elements

Observations in inclusions of meteorites

^{48}Ca : -the last stable neutron rich nucleus in Ca chain
-two neutrons richer than ^{46}Ca which is also stable

-on Earth $^{48}\text{Ca}/^{46}\text{Ca} \sim 50 \gg 1$

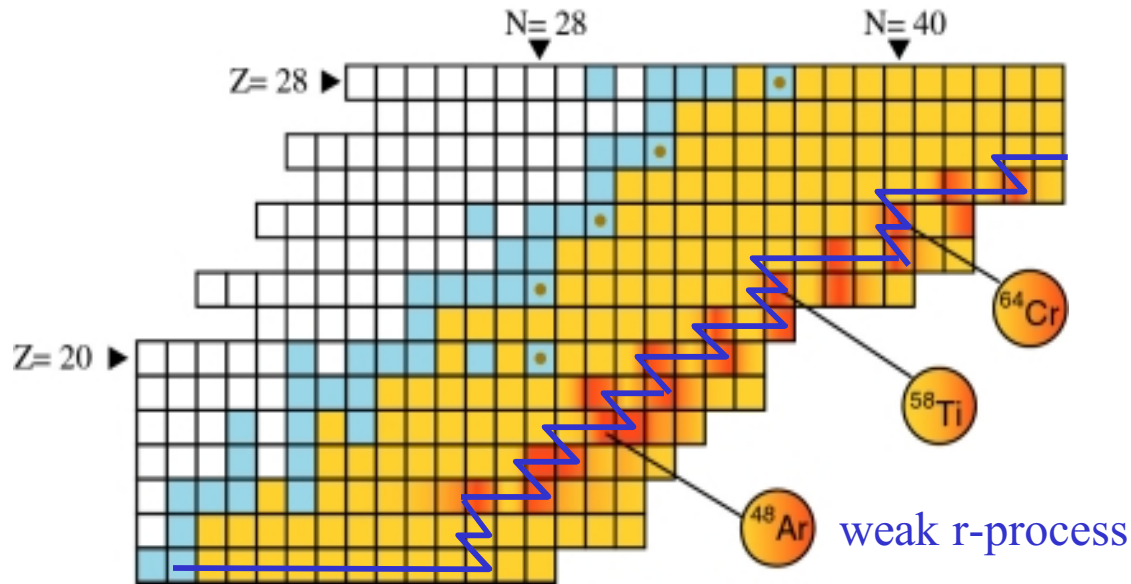
BUT :

-EK-1-4-1 $^{48}\text{Ca}/^{46}\text{Ca} \sim 250$

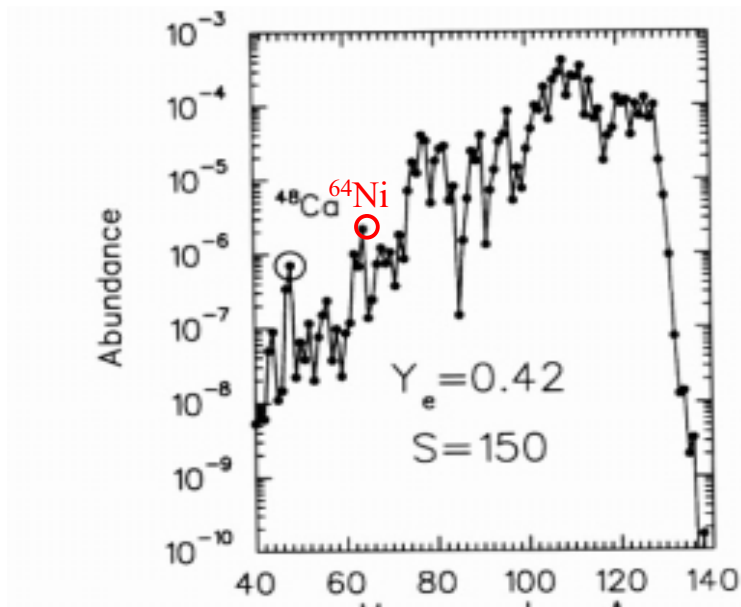
And also **correlated isotopic anomalies** in neutron-rich nuclei ^{50}Ti - ^{54}Cr - ^{58}Fe - ^{64}Ni

Could be explain by a “weak” r-process

The weak r-process : Production of $A < 130$ elements



- Explosive environment
 - Involves neutron rich nuclei
- Neutron captures
 β decays

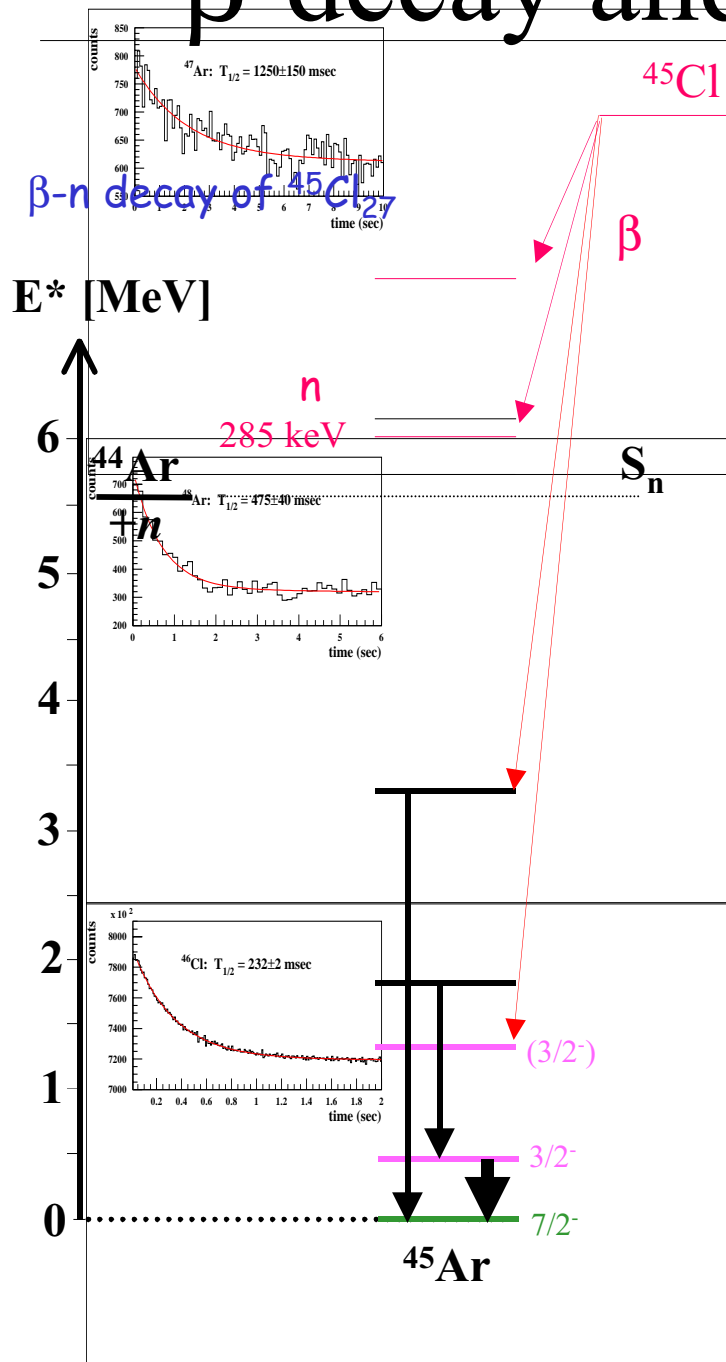


- Production of nuclei with $A < 130$
- Overproduction of ^{48}Ca & ^{64}Ni

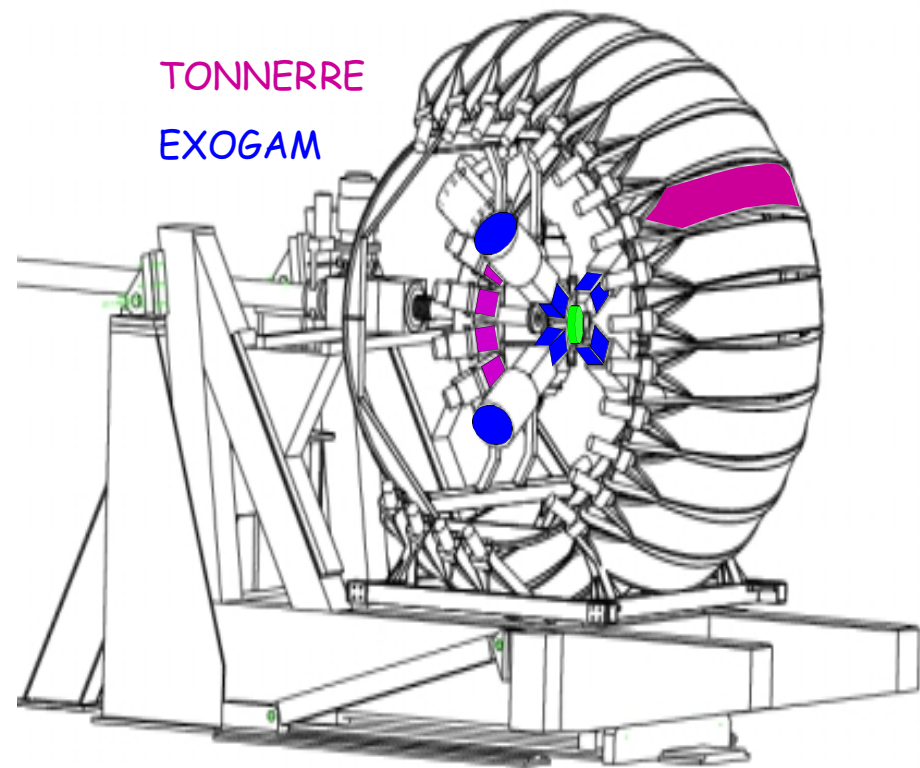
Mass Number A

K.L. Kratz et al. Proc. Della Societa Astronomica Italiana 2000

β -decay and spectroscopy studies



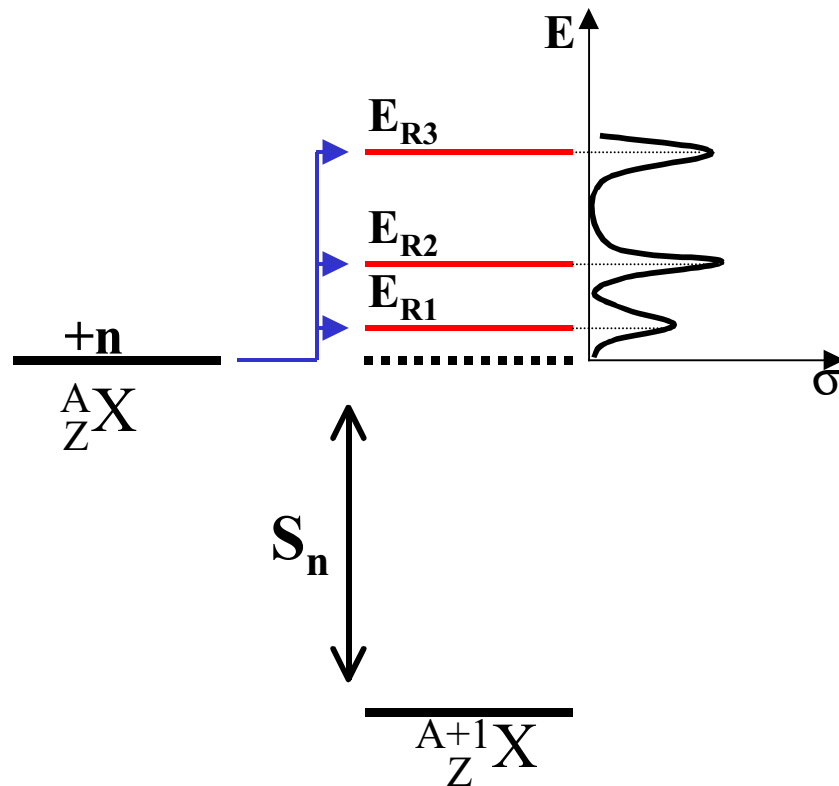
- Some states could contribute to the direct capture



S. Grévy, private comm., LPC Caen

Neutron capture mechanisms

Compound nucleus : resonant captures



-If **few** resonances above the threshold :
resonant capture
=>application of Breit-Wigner formalism

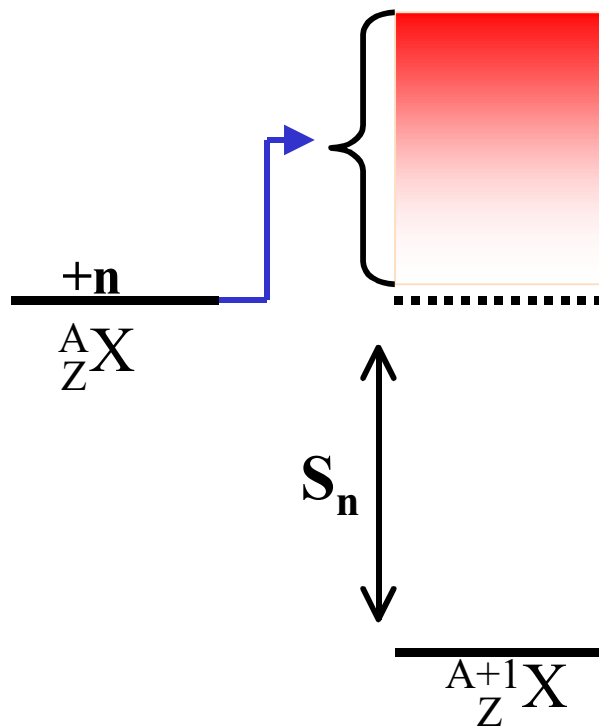
-Captures only possible for **given** neutron
energies : $E_n = S_n + E_R$

-**Strong** variation of the cross-section
with the neutron energy depending on
quantum numbers of the final states

Needs structural informations of unbound states in the final nucleus

Neutron capture mechanisms

Compound nucleus : statistical treatment



-**High** level density above the threshold :
overlap between resonances
=> application of Hauser-Feshbach formalism

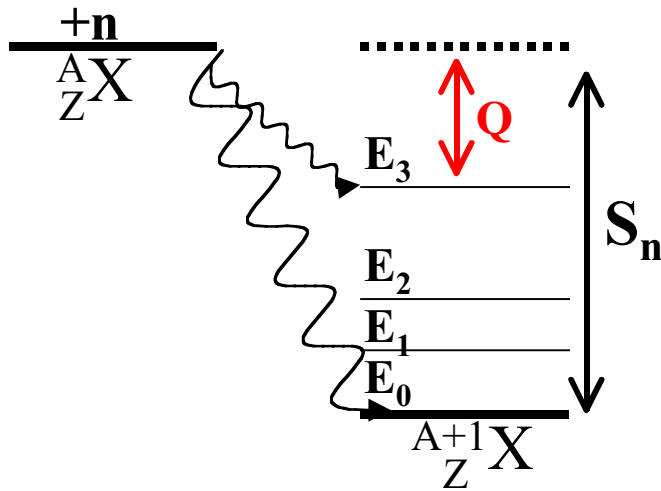
-Captures possible for **all** neutron energies above the threshold

-**Smooth** variation of the cross-section with the neutron energy

Determination of the cross section via statistical models : applicability (structure, S_n)?

Neutron capture mechanisms

Direct capture



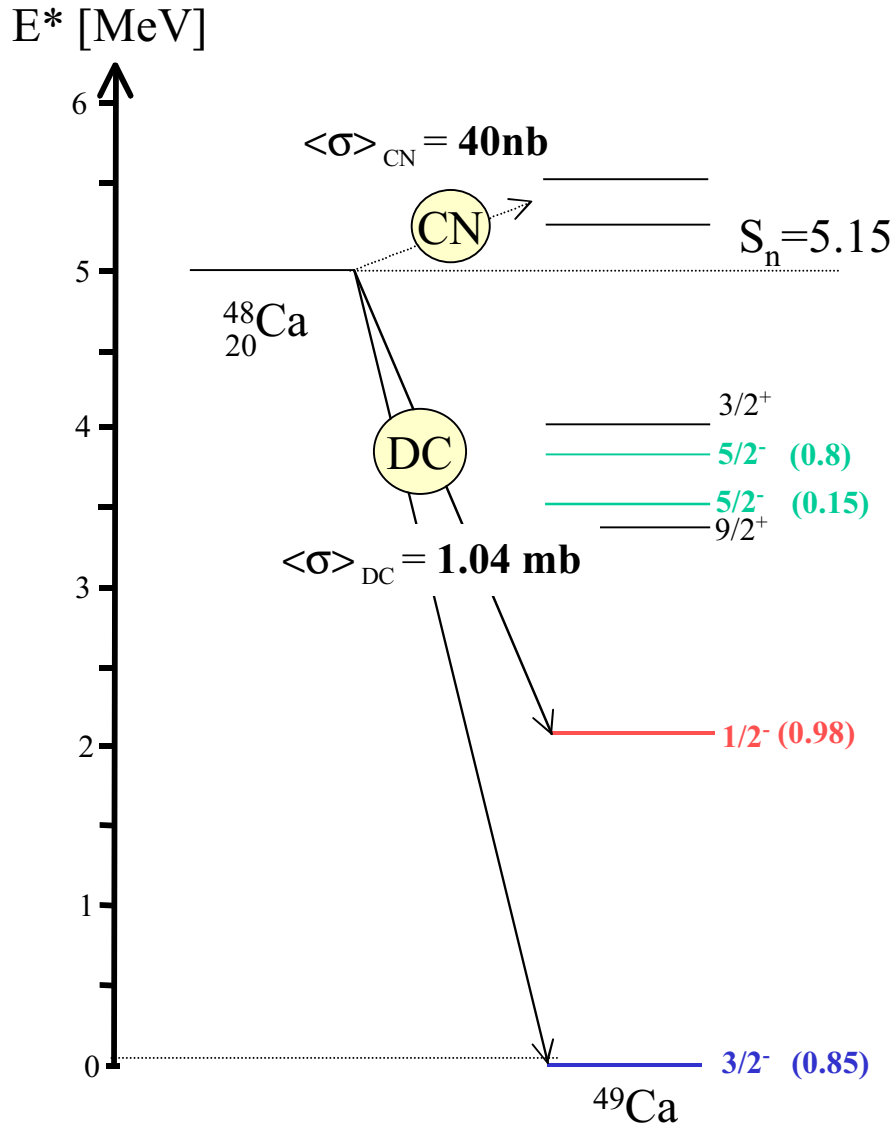
-Captures on bound levels in final nucleus

-Captures possible for **all** neutron energies

-**Smooth** variation of the cross-section with the neutron energy depending on **l** , **Q** and **S** of final state

Needs structural informations of low energy bound states : spectroscopic factors accessible by transfer reactions

Neutron capture on ^{46}Ar



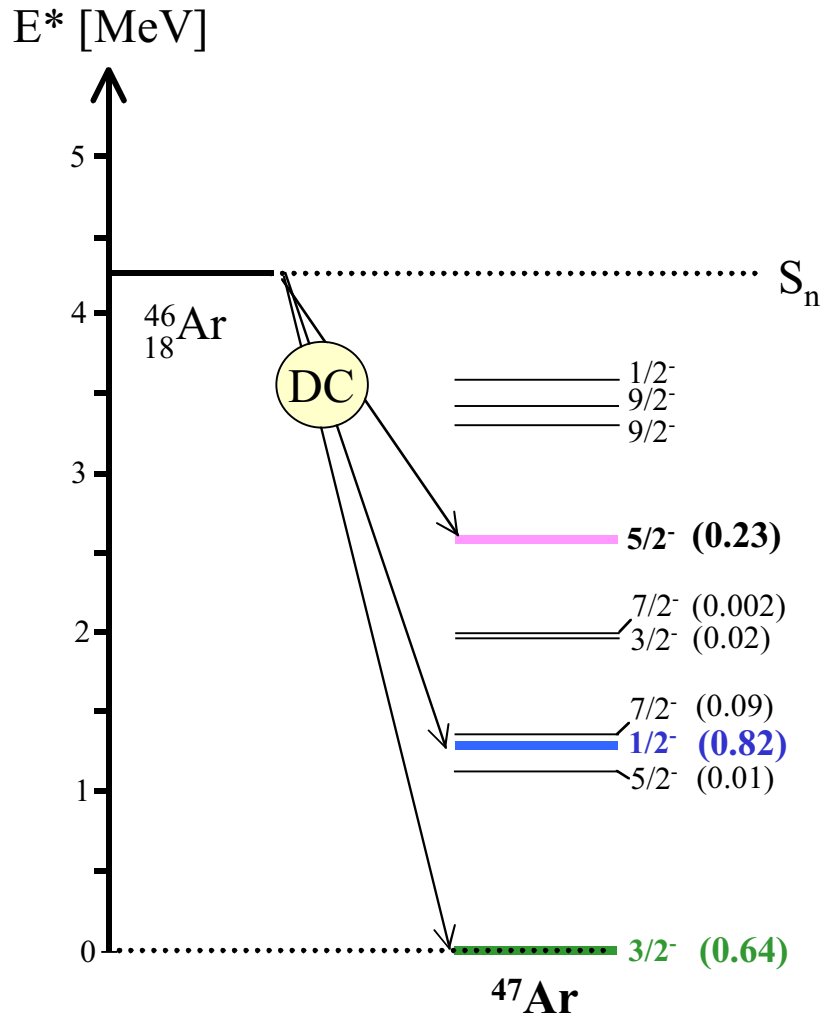
*E. Kraussmann et al. PRC 53 (1996) 469.
R. Abegg et al. NPA 303 (1978) 121.*

- **N=28** magic number
- **Low** neutron separation energy

↳ The level density around S_n assumed to be low : necessity to consider direct capture

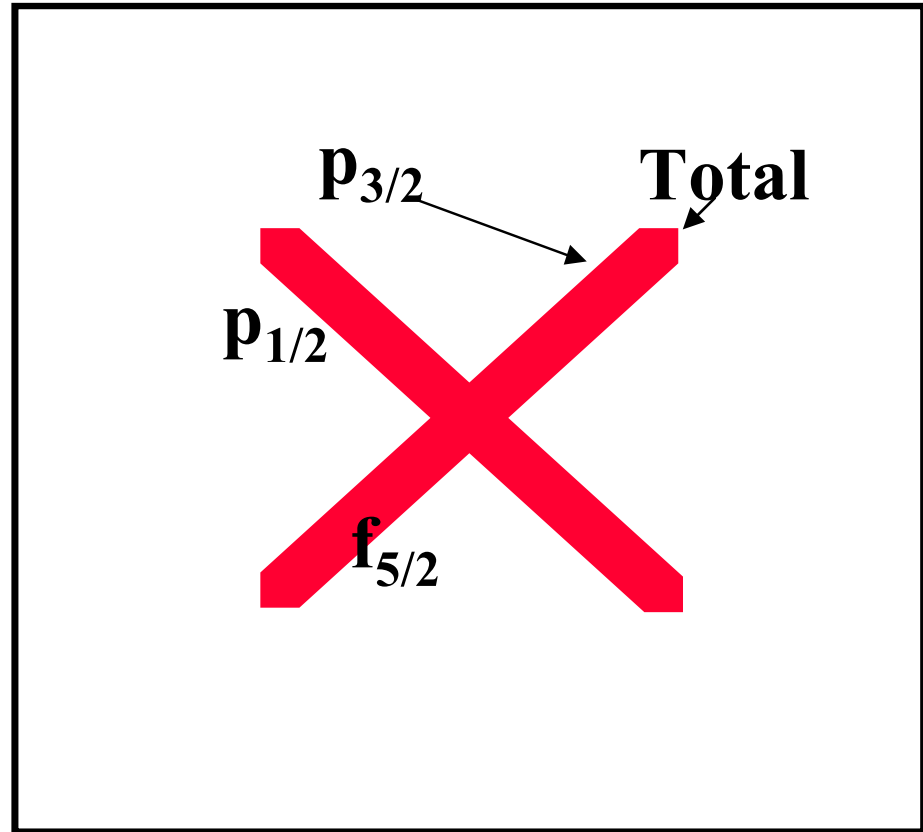
↳ 95% of the cross section is DC

Neutron capture on ^{46}Ar



Shell Model results

ANTOINE



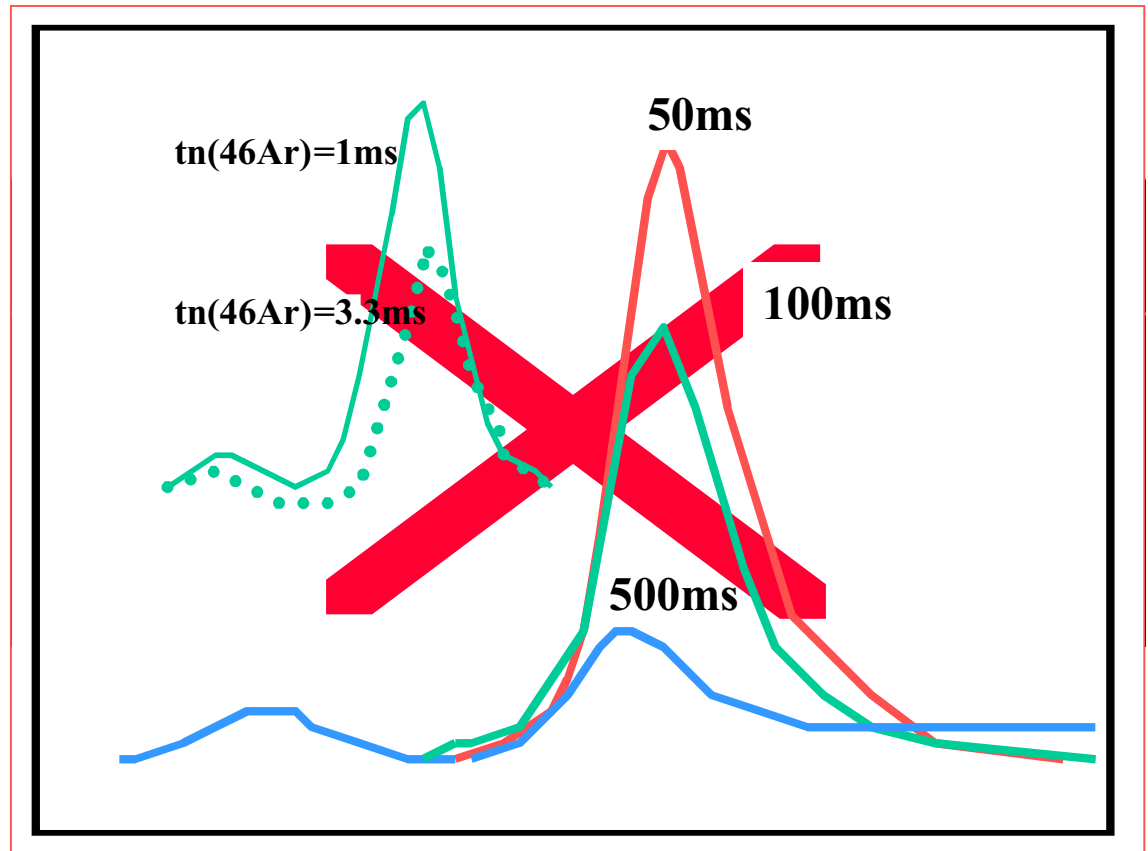
Nuclear structure of ^{47}Ar enhances neutron capture on ^{46}Ar which is a N=28 closed shell nucleus

Weak r-process network simulation

Including : β -decays, neutron captures for nuclei from stability to neutron drip-line

Require to solve more than 1000 coupled differential equations.

Not yet time dependent :
 $T_9 = \text{cst}$, $d_n = \text{cst}$



Conclusion & outlooks

- **Weak r-process** could be responsible for production of **low mass neutron rich nuclei**
- Needs of **β half-lives**, **neutron capture** cross-sections for prediction of the weak r-process final abundances
- Neutron-rich **magic** nuclei \Rightarrow mainly **direct capture** mechanism
 \Rightarrow **low energy structure** in final nucleus
- **Complementarity** of **β -decay- γ -spectroscopy** and **transfer reaction**
 - \Rightarrow **energy of low lying states**
 - \Rightarrow **spins & spectroscopic factors**

Coming soon :

Transfer reaction at GANIL, in June : $^{44,46}\text{Ar}(d,p)$

\Rightarrow energy, spins, spectro. fac. (to be **compared** with already **measured** levels in ^{45}Ar)

\Rightarrow **size of the N=28 gap**

Collaboration

Isotopic anomalies :

O. Sorlin, L. Gaudefroy (Institut de Physique nucléaire)

K.L. KRATZ (Institut für Kerchemie; Mainz university)

T. Rauscher (Departement für Physik und astronomy; Basel university)

Transfer reaction in Ar isotopes :

Institut de Physique nucléaire

CEA Saclay

GANIL

LPC Caen

Institute of Nuclear Research (Hungary)

FLNR-Dubna (Russia)

NBI (University of Copenhagen)

Direct neutron capture : TEDCA

$$\frac{f\sigma}{f\Omega} \propto S \left| \int \Phi_f \Theta_{EM} \Phi_i d\vec{r} \right|^2$$

Φ_f, Φ_i : coming from resolution of Schrödinger equation with a potential obtained by double folding

Θ_{EM} : E1, E2, M1 electromagnetic operator

$$V = \lambda \int \rho_n(\vec{r}_n) \rho_A(\vec{r}_A) V_{eff}(E, \rho_n, \rho_A) d\vec{r}_n d\vec{r}_A$$

V_{eff} : DDM3Y interaction

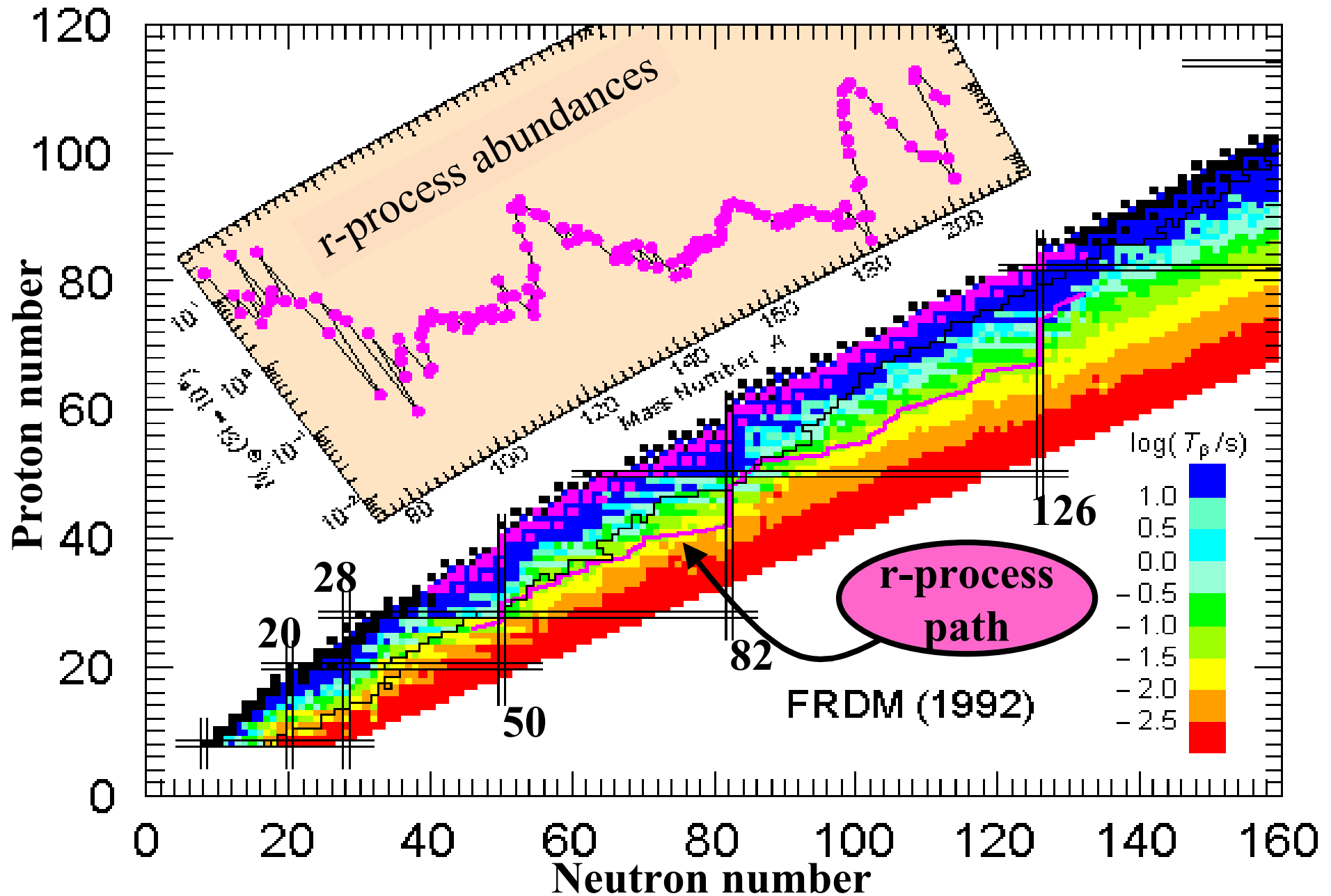
Neutron density ρ_n : δ function

Target nucleus density ρ_A : mean field theory (HFB)

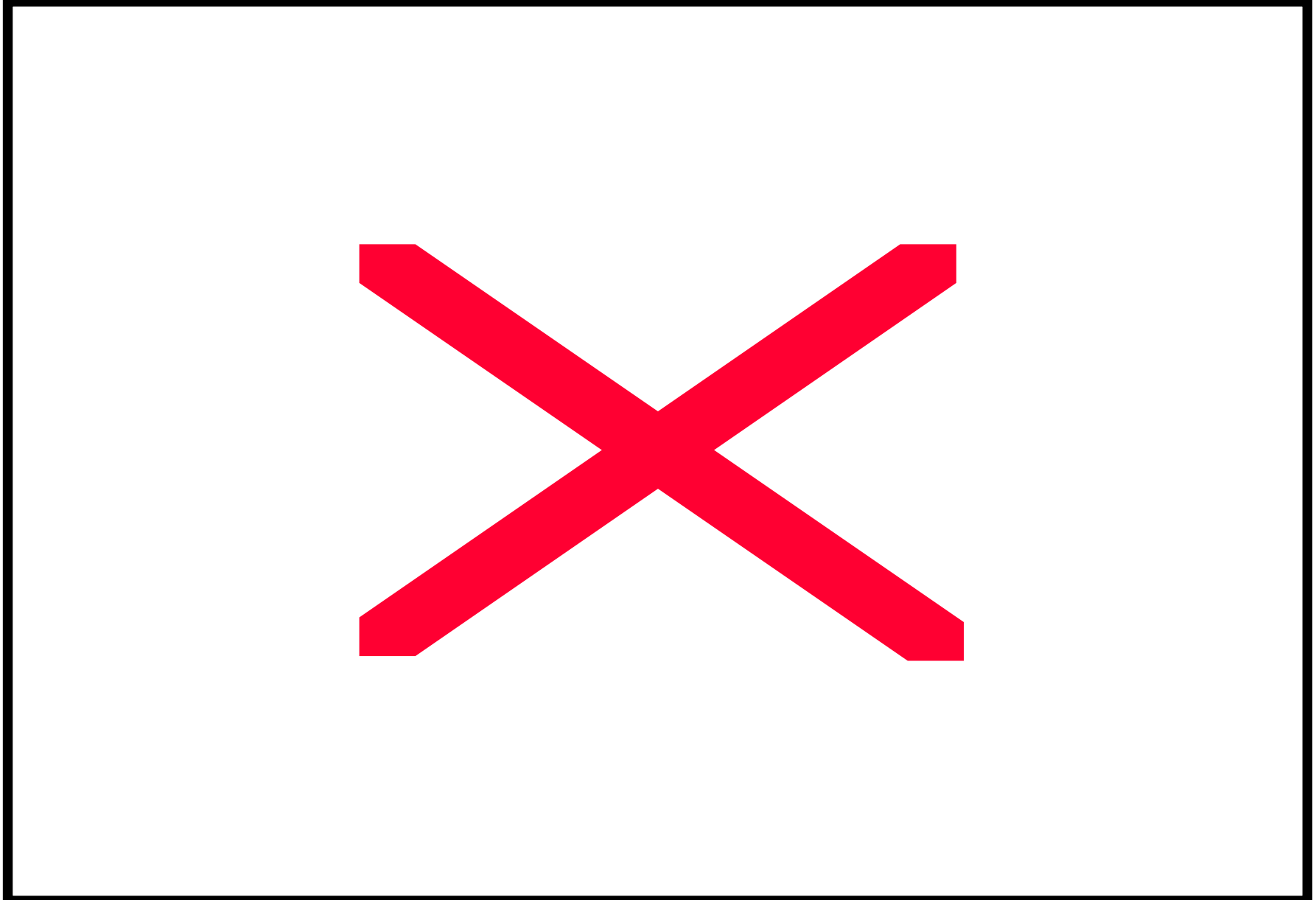
Free parameter λ :

- Entrance channel : fixed to reproduced neutron scattering
- Exit channel : fixed to reproduce neutron separation energy

Neutron captures in the (weak) r-process

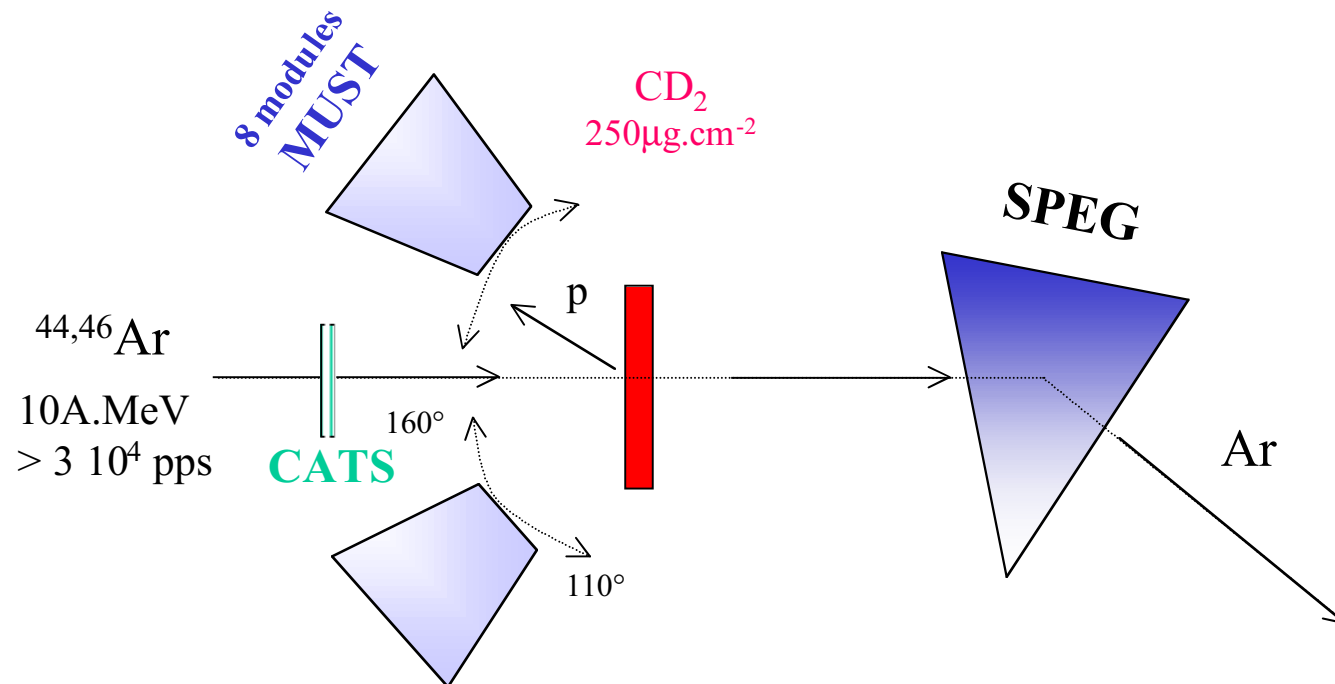


Neutron captures in the (weak) r-process



Neutron capture on $^{44,46}\text{Ar}$ (Experiment)

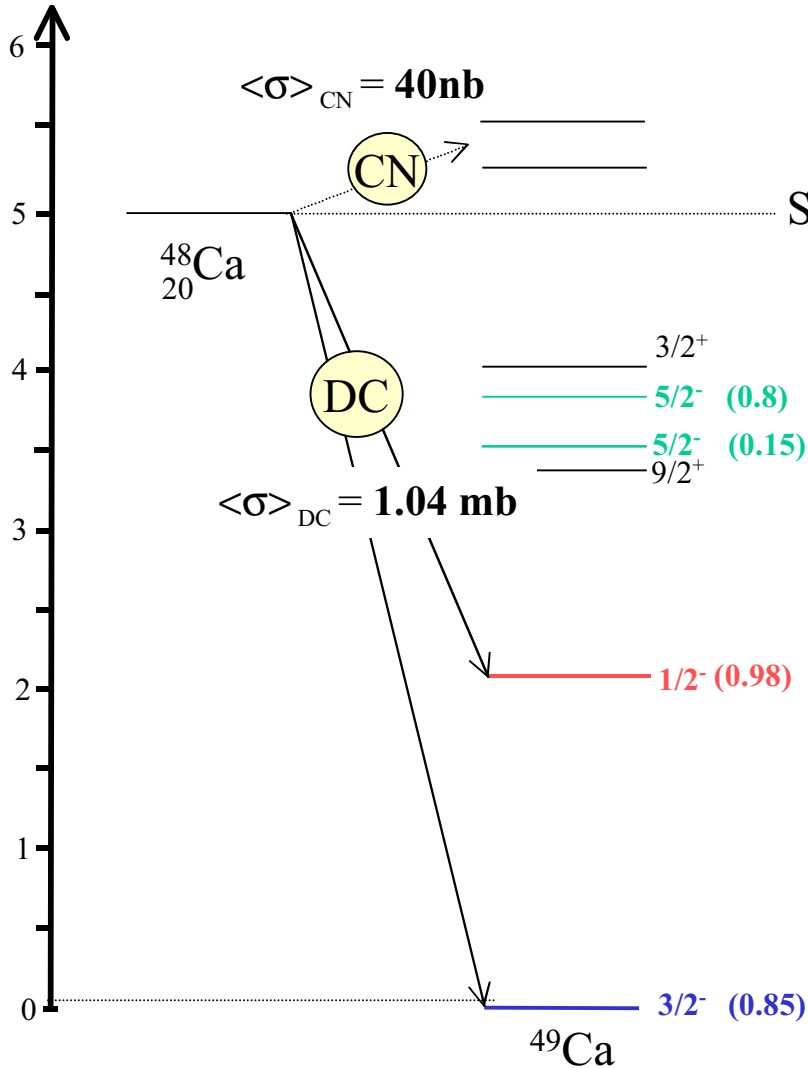
- Experiment in June at GANIL



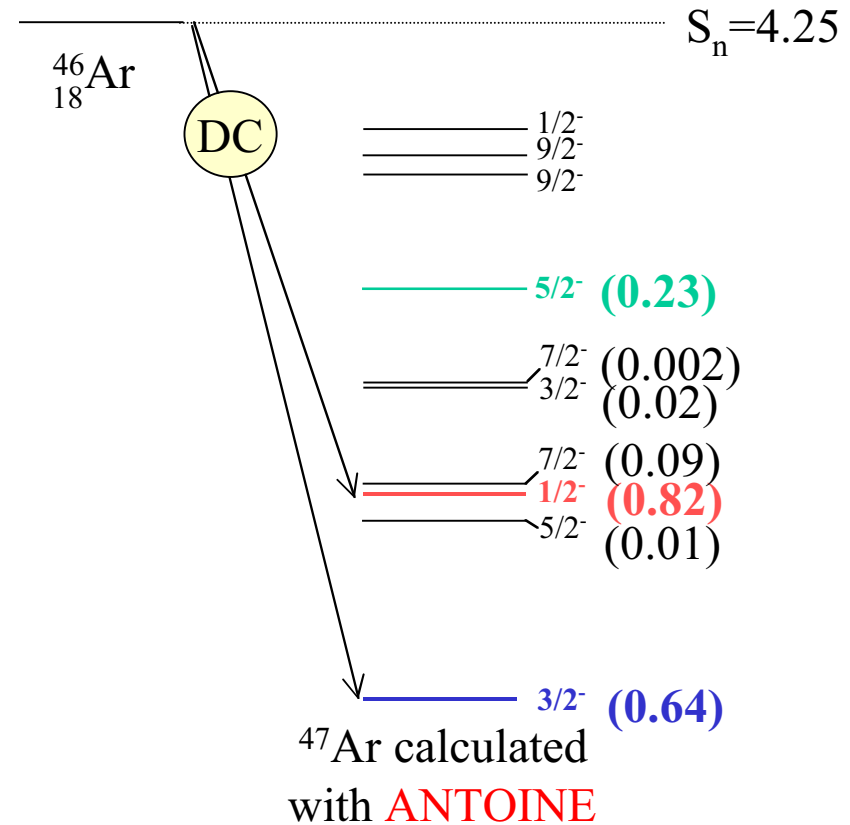
Study of the $p_{3/2}$, $p_{1/2}$ and $f_{5/2}$ states: energy, spec. factor, spin

Neutron capture on $^{44,46}\text{Ar}$

E^* [MeV]



E. Krausmann et al. PRC 53 (1996) 469.
R. Abegg et al. NPA 303 (1978) 121.



Similar behaviour expected in ^{46}Ar

95% of the cross section is DC